

Modeling Of Tokamak Plasmas

J.D. Callen, University of Wisconsin, Madison, WI 53706-1609 USA

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*Questions To Be Addressed:*¹

- 1) What effects need to be modeled for ITER plasmas?
- 2) What equations need to be solved for integrated simulations?
- 3) How can MHD, kinetic and transport models be self-consistent?

Outline:

- Modeling ITER plasmas involves disparate \vec{x} , t scales and issues.
- Comprehensive modeling of plasmas involves some key elements:
plasma kinetic equation with Fokker-Planck collision operator and sources,
fluid equations when $t > 1/\nu$, Chapman-Enskog (C-E) kinetic equation,
extended MHD (ideal MHD for $t < 1/\nu$, neoclassical MHD for $t > 1/\nu$),
collisional-, fluctuation- and 3-D- induced fluxes via C-E kinetic equation,
and comprehensive plasma transport equations for n_e , Ω_{tor} , p_e , p_i and ψ_p .
- Need modular approach for integrated modeling of ITER plasmas.

¹J.D. Callen, CEMRACS 2014 “Fluid and transport modeling of plasmas” lectures available via <http://homepages.cae.wisc.edu/~callen/plasmas>.

This Evolving Study Has A Long-Term Objective

- Develop a strategic vision for how the major plasma models (MHD, kinetics, transport) that operate on a hierarchy of sequentially longer time scales can and should be combined into self-consistent, comprehensive, integrated simulations of plasmas in present tokamaks to develop a “predictive capability” for ITER:
 - what physical effects should be included?,
 - what equations need to be solved?,
 - how can they be coupled and made self-consistent?
- This presentation will:
 - discuss logic of combining MHD, kinetics, transport approaches,
 - highlight effects of diamagnetic-level flows and small 3-D fields,
 - focus on descriptions inside separatrix, but can be generalized.

Characteristic Length And Time Scales In Plasmas Span Many Orders of Magnitude (ITER and ICF)

- Projected parameters for ITER where $B_t = 5.6$ T, $T_e \sim T_i \sim 10$ keV, $n_e \sim 10^{20} \text{ m}^{-3}$, and major/mid-plane minor radius $\simeq 6 \text{ m}/2 \text{ m}$ are:

Length Scales

minimum impact distance	b_{\min}^{qm}	10^{-12} m
mean particle spacing	$n_e^{-1/3}$	2×10^{-7} m
Debye shielding length	λ_{De}	7×10^{-5} m
deuteron gyroradius	ρ_D	3×10^{-3} m
average minor radius	\bar{a}	3 m
collision length	λ_e	1.2×10^4 m

Time Scales

plasma period	$1/\omega_p$	2×10^{-12} s
deuteron gyroperiod	$1/\omega_{cD}$	3×10^{-9} s
Alfvén period	\bar{a}/c_A	5×10^{-7} s
sound wave period	\bar{a}/c_S	4×10^{-6} s
electron collision time	$1/\nu_e$	2×10^{-4} s
deuteron collision time	$1/\nu_D$	3×10^{-2} s
energy confinement time	τ_E	6 s
fusion collision time	$1/\nu_{\text{fus}}$	200 s
magnetic field diffusion	τ_R	1000 s

Dimensionless Parameters

	<u>formula</u>	<u>ICF</u>	<u>ITER</u>
number of electrons in a Debye cube	$n_e \lambda_{De}^3$	3.5×10^3	4×10^7
electron collision to plasma frequency	ν_e/ω_p	4×10^{-5}	10^{-8}
ρ_*: deuteron gyroradius to average minor radius	ρ_D/\bar{a}		10^{-3}
collision length to relevant length	λ_e/L	$\lambda_e/\bar{a} \sim 0.4$	$\lambda_e/2\pi R = 320$

Integrated Simulations Should Include Many Effects

- Current ramp-up from breakdown to current flattop, smooth turnoff: tearing-limited resistive evolution of $\vec{J}(\rho, t)$; NBI torque to avoid locked modes.
- MHD stability boundaries and tearing-induced topology changes: low q_{95} and high β_N stability limits; slowly growing tearing-induced islands.
- Use of 3-D fields to control plasma rotation and H-mode pedestals: NTV control of toroidal plasma rotation; RMPs to mitigate/suppress ELMs.
- Heating, momentum, particle and poloidal flux sources and sinks: core NBI, ICRF, EC; edge neutral recycling, charge-exchange, radiation etc.
- Transport evolution and control of n_e , $\Omega_{\text{tor}}(E_\rho)$, T_e , T_i , ψ_p profiles: equil. & stiff ones, internal & H-mode transport barriers, hybrid flux pumping.

All these effects need to be validated and modeled self-consistently to develop a predictive capability for ITER plasmas.

MHD, Kinetics, Transport Are Usually Treated Separately

- **MHD models** can provide

axisymmetric equilibrium — force balances, $P = P(\psi_p)$, $\vec{B}_0 = I\vec{\nabla}\zeta + \vec{\nabla}\zeta \times \vec{\nabla}\psi_p$;

ideal MHD stability constraints — minimum $q(\psi_p)$, maximum β ;

neoclassical equilibrium flows — || Ohm's law with bootstrap current, $V_{pol i}$;

magnetic reconnection of field lines at rational surfaces in dissipative MHD

— tearing instabilities, penetration of external $\delta\vec{B}$ s, locked modes; and

plasma amplification of external 3-D $\delta\vec{B}$'s that couple to least stable modes.

- **Kinetic models** can provide transport fluxes in $\vec{B}_0 + \delta\vec{B}$ due to

Coulomb collisional effects — classical, neoclassical, paleoclassical;

fluctuation-induced microturbulence — ITG, TEM, KBM etc.; and

small 3-D magnetic fields — error fields, ripple, NTV, RMPs, flutter, islands.

- **Plasma transport models** are based on axisym. \vec{B}_0 surfaces and use

fluxes from kinetics plus sources and sinks in transport evolution equations

to determine the ρ, t behavior of plasma parameters on transport time scale.

Comprehensive integrated modeling of tokamak plasmas needs a self-consistent theoretical framework for combining these models.

Fundamental Equation Is Plasma Kinetic Equation (PKE)

- **PKE is**
$$\frac{df_s(\vec{x}, \vec{v}, t)}{dt} \equiv \frac{\partial f_s}{\partial t} + \vec{v} \cdot \frac{\partial f_s}{\partial \vec{x}} + \frac{\vec{F}_s}{m_s} \cdot \frac{\partial f_s}{\partial \vec{v}} = \mathcal{C}\{f_s\} + \mathcal{S}\{f_s\},$$
 in which

$f_s \equiv f_s(\vec{x}, \vec{v}, t)$ is the distribution function in 6-D (\vec{x}, \vec{v}) phase space,

$\vec{F}_s \equiv q_s(\vec{E} + \vec{v} \times \vec{B})$ is the Lorentz force on a charged particle,

$\mathcal{C}\{f_s\}$ is the Fokker-Planck Coulomb collision operator which is local in \vec{x} , and

$\mathcal{S}\{f_s\}$ is source operator that represents heating, current-drive, c-x etc. effects.

- While the PKE is fundamental, the 3-D (\vec{x}) fluid moment equations for n_s , \vec{V}_s , p_s are

feasible, appropriate, useful and needed,

exact to extent that relevant Chapman-Enskog kinetic equation (CEKE, p 12,13)

can be solved for kinetic distortion $F_s \equiv f_s - f_{\text{Max } s}(\vec{x}, \vec{v}, t)$,

which then yields the needed closure and collisional moments, and

provide the basis for both extended MHD and plasma transport equations.

Species s Fluid Moment Equations Are Useful, Needed

- The $\int d^3v (1, m_s \vec{v}, m_s v^2/2)$ moments of the plasma kinetic equation (PKE) yield the species s complete fluid moment equations:

$$\text{density} \quad (\partial/\partial t + \vec{V}_s \cdot \vec{\nabla}) n_s = -n_s \vec{\nabla} \cdot \vec{V}_s + S_{ns},$$

$$\text{momentum} \quad m_s n_s (\partial/\partial t + \vec{V}_s \cdot \vec{\nabla}) \vec{V}_s = n_s q_s (\vec{E} + \vec{V}_s \times \vec{B}) - \vec{\nabla} p_s - \vec{\nabla} \cdot \overleftrightarrow{\pi}_s + \vec{R}_s + \vec{S}_{\vec{p}s},$$

$$\text{energy} \quad \frac{3}{2} (\partial/\partial t + \vec{V}_s \cdot \vec{\nabla}) p_s = -\frac{5}{2} p_s \vec{\nabla} \cdot \vec{V}_s + p_s \dot{s}_{Ms}, \quad \text{or,}$$

$$\text{entropy} \quad (\partial/\partial t + \vec{V}_s \cdot \vec{\nabla}) s_{Ms} = \dot{s}_{Ms} \equiv (-\vec{\nabla} \cdot \vec{q}_s - \overleftrightarrow{\pi}_s : \vec{\nabla} \vec{V}_s + Q_s + S_{\mathcal{E}s})/p_s,$$

in which the species s isotropic-Maxwellian-based collisional entropy is

$$s_{Ms}(\vec{x}, t) \equiv -\frac{1}{n} \int d^3v f_{iMs} \ln f_{iMs} = \frac{3}{2} \ln \left(\frac{p_s}{n^{5/3}} \right) + C, \quad \text{collisional entropy.}$$

- But equations are incomplete until these “closures” are specified:

stress $\overleftrightarrow{\pi}_s$, heat flux \vec{q}_s ; collisional friction force \vec{R}_s , energy exchange Q_s .

Modeling Tokamaks Employs Some Key Approximations

- The fundamental approximation used in modeling tokamak plasmas is that the gyroradii ($\rho_i \equiv v_{Ti}/\omega_{ci}$) of the charged ions in the magnetic field \vec{B} are small compared to macroscopic gradient scale lengths $L \sim 1/|\vec{\nabla} \ln(B, n, T)|$, i.e., $\rho_* \equiv \rho_i/L \ll 1$. This facilitates analyses on a hierarchy of ever longer time scales:
 - ρ_*^0 : parallel (to \vec{B}) guiding center motion, Alfvén and sound waves ($\sim \mu\text{s}$),
 - ρ_*^1 : particle drifts across the magnetic field, fluid descriptions, collisional effects on species flows in flux surfaces, magnetic reconnection ($\gtrsim \text{ms}$), and finally
 - ρ_*^2 : transport of plasma across magnetic field lines and flux surfaces due to collisions, radially localized microturbulence and 3-D field effects ($\gtrsim \text{s}$).
- Sequential ρ_* orders cause momentum balance component effects:
 - $\rho_*^0, \vec{\nabla} \psi_p \cdot \Rightarrow$ MHD comp. Alfvén waves enforce ($\sim 0.4 \mu\text{s}$) radial force balances,
 - $\rho_*^1, \vec{B}_0 \cdot \Rightarrow$ drifts, neoclassical MHD flows within ψ_p surfaces ($e 0.2 \text{ ms}, i 30 \text{ ms}$),
 - $\rho_*^2, \vec{e}_\zeta \cdot \Rightarrow$ toroidal torques on plasma cause radial transport fluxes ($\tau_E \sim 6 \text{ s}$).

Extended MHD Model Includes Ideal MHD And The Dissipative Effects Of Closure And Collisional Moments

- The extended MHD equations for a magnetized plasma are obtained by summing the fluid moment equations over species. Together with equations for the magnetic field they are

Extended MHD plasma description (for ideal MHD \vec{R}_e , $\overleftrightarrow{\Pi}$, $\overleftrightarrow{\pi}_e$, $\sum_s \dot{s}_{Ms} \rightarrow 0$):

mass density $(\partial/\partial t + \vec{V} \cdot \vec{\nabla}) \rho_m = -\rho_m \vec{\nabla} \cdot \vec{V},$

charge continuity $\vec{\nabla} \cdot \vec{J} = 0,$

momentum $\rho_m(\partial/\partial t + \vec{V} \cdot \vec{\nabla})\vec{V} = \vec{J} \times \vec{B} - \vec{\nabla} P - \vec{\nabla} \cdot \overleftrightarrow{\Pi},$

Ohm's law $\vec{E} = -\vec{V} \times \vec{B} + \vec{R}_e/n_e e + (\vec{J} \times \vec{B} - \vec{\nabla} p_e - \vec{\nabla} \cdot \overleftrightarrow{\pi}_e)/n_e e,$

equation of state $(\partial/\partial t + \vec{V} \cdot \vec{\nabla}) \ln(P/\rho_m^{5/3}) = \sum_s \dot{s}_{Ms}.$

Maxwell equations for extended MHD (no Gauss' law, \vec{E} from Ohm's law):

Faraday's law $\partial \vec{B} / \partial t = -\vec{\nabla} \times \vec{E},$

no magnetic monopoles $\vec{\nabla} \cdot \vec{B} = 0,$

nonrelativistic Ampere's law $\vec{J} = \vec{\nabla} \times \vec{B} / \mu_0.$

Extended MHD Equations Produce Many Effects

- Ideal MHD in a lowest order axisymmetric tokamak provides equilibrium radial force balance that is enforced by compressional Alfvén waves on $\bar{a}/c_A \sim 0.5 \mu\text{s}$ time scale and yields Grad-Shafranov equation for ψ_p , frozen flux theorem (\vec{B} advects with \vec{V}_\perp) due to $\vec{E} + \vec{V} \times \vec{B} = \vec{0}$ Ohm's law, ideal MHD constraints for kink (min q), $\vec{\nabla}P$ -driven (max β) stability, and amplification of driven 3-D perturbations that couple to least stable modes.
- Two-fluid MHD causes a diamagnetic-level flow constraint because $\vec{\nabla}\psi_p \cdot$ its Ohm's law $\vec{E} + \vec{V} \times \vec{B} = \vec{\nabla}p_i/n_iq_i$ yields relation between flows in surface

$$\Omega_{\text{tor}} \equiv \vec{V}_i \cdot \vec{\nabla}\zeta = - \left(\frac{\partial\Phi_0}{\partial\psi_p} + \frac{1}{n_iq_i} \frac{\partial p_i}{\partial\psi_p} \right) + q \vec{V}_i \cdot \vec{\nabla}\theta \implies V_{\text{tor}} \simeq \frac{E_\rho}{B_p} - \frac{1}{n_iq_i B_p} \frac{dp_i}{d\rho} + \frac{B_t}{B_p} V_{\text{pol}i}.$$
- Extended MHD with collision-based closures for $t > 1/\nu$ leads to reconnection of field lines in dissipative singular layers at rational surfaces, classical and neoclassical tearing-type instabilities \implies magnetic islands, FSA || neo Ohm's law $\langle \vec{B}_0 \cdot \vec{E}^A \rangle = \eta_{\parallel}^{\text{nc}} \left[\langle \vec{B}_0 \cdot \vec{J} \rangle - \langle \vec{B}_0 \cdot \vec{J}_{\text{drives}} \rangle \right]$, for $t > 0.2$ ms, poloidal ion flow damped to $V_{\text{pol}i}^{\text{nc}} \simeq (c_p/q_i B) (dT_i/d\rho)$, $c_p \simeq 1.17$, for $t > 30$ ms, 3-D \vec{B} field modifications and their effects — error field locking, NTV, RMPs.

Extended MHD Constraints Should Be Used In Kinetics

- Lowest order constraints for MHD stable plasmas:

ϱ_*^0 : axisymmetric magnetic field $\vec{B}_0 \equiv I\vec{\nabla}\zeta + \vec{\nabla}\zeta \times \vec{\nabla}\psi_p = \vec{\nabla}\psi_p \times \vec{\nabla}(q\theta - \zeta)$, with poloidal magnetic flux surfaces ψ_p determined from Grad-Shafranov equation,

ϱ_*^1 : a consistency relation between the flows of $V_{\text{tor}} \simeq \frac{E_\rho}{B_p} - \frac{1}{n_i q_i B_p} \frac{dp_i}{d\rho} + \frac{B_t}{B_p} V_{\text{pol}i}$ — which does not determine the electric field E_ρ or plasma toroidal rotation frequency $\Omega_{\text{tor}} \simeq V_{\text{tor}}/R$, but provides a relation between them.

- First order (ϱ_*^1) extended MHD “equilibrium effects” for tearing-mode stable plasmas that could be used to speed convergence:

Equilibrium Maxwellian distribution could include the following ϱ_* flows: parallel electron flow that yields the FSA || neoclassical Ohm’s law, and the neoclassical poloidal ion flow $V_{\text{pol}i}^{\text{nc}}$.

A magnetic field structure $\vec{B} = \vec{B}_0 + \delta\vec{B}$ in which the ϱ_* perturbations $\delta\vec{B}$ are nonzero in dissipative singular layers near rational surfaces, and are driven in plasma outside them by small externally-imposed 3-D fields.

Use Chapman-Enskog Approach For Kinetic Equation

- Chapman-Enskog Ansatz posits species distribution function has two parts — “dynamic” space- and time-dependent Maxwellian f_M with parameters $n(\vec{x}, t)$, $\vec{V}(\vec{x}, t)$, $T(\vec{x}, t)$ plus kinetic distortion F :

$f(\vec{x}, \vec{v}, t) = f_M(\vec{x}, \vec{v}, t) + F(\vec{x}, \vec{v}, t)$, in which dynamic Maxwellian is

$$f_M(\vec{x}, \vec{v}, t) = f_M[n(\vec{x}, t), \vec{V}(\vec{x}, t), T(\vec{x}, t), \vec{v}] = \frac{n(\vec{x}, t) e^{-m [\vec{v} - \vec{V}(\vec{x}, t)]^2 / 2T(\vec{x}, t)}}{[2\pi T(\vec{x}, t)/m]^{3/2}}.$$

- Substituting this Ansatz into the plasma kinetic equation yields

$$\boxed{\frac{dF}{dt} - \mathcal{C}\{f\} - \mathcal{S}\{f\} = -\frac{df_M}{dt}}, \quad \text{general Chapman-Enskog kinetic equation,}$$

$$\begin{aligned} \frac{df_M}{dt} = & f_M \left[-\frac{1}{p} \vec{v}_r \cdot \left(nq [\vec{E} + \vec{V} \times \vec{B}] - \vec{\nabla} p - mn \left(\frac{\partial}{\partial t} + \vec{V} \cdot \vec{\nabla} \right) \vec{V} \right) \right. && \text{forces} \\ & + \left(\frac{mv_r^2}{2T} - \frac{5}{2} \right) \frac{1}{T} \vec{v}_r \cdot \vec{\nabla} T && T \text{ gradient} \\ & + \frac{m}{T} \left(\vec{v}_r \vec{v}_r - \frac{v_r^2}{3} \mathbf{I} \right) : \mathbf{W}, \quad \mathbf{W} \equiv \frac{1}{2} [\vec{\nabla} \vec{V} + (\vec{\nabla} \vec{V})^T - \frac{2}{3} \mathbf{I} \vec{\nabla} \cdot \vec{V}] && \text{rate of strain} \\ & + \frac{S_n}{n} + \left(\frac{mv_r^2}{2T} - \frac{3}{2} \right) \left(\frac{2 \dot{s}_M}{3p} - \frac{S_n}{n} \right) \left. \right]. && \text{xport sources} \end{aligned}$$

Chapman-Enskog Approach Is Useful And Important

- **Chapman-Enskog kinetic equation (CEKE)** on preceding page is still exact since no approximations or truncations have been made.

- Lowest order drift-kinetic CEKE obtained by using momentum equation from p 7 and gyro-averaging yields ($\hat{\mathbf{b}} \equiv \vec{B}/B_0$)^{2,3,4}

$$\begin{aligned} & \frac{\partial \bar{F}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \vec{v}_d) \cdot \frac{\partial \bar{F}}{\partial \vec{x}} - \overline{\mathcal{C}\{f_M + F\}} - \overline{\mathcal{S}\{f_M + F\}} \\ & = \bar{f}_M \left[\left(\frac{mv^2}{2T} - \frac{5}{2} \right) v_{\parallel} \hat{\mathbf{b}} \cdot \vec{\nabla} \ln T + \frac{m}{T} \left(v_{\parallel}^2 - \frac{v_{\perp}^2}{2} \right) [\hat{\mathbf{b}} \cdot (\hat{\mathbf{b}} \cdot \vec{\nabla}) \vec{V} - \vec{\nabla} \cdot \vec{V} / 3] - \frac{v_{\parallel}}{p} \hat{\mathbf{b}} \cdot [\vec{\nabla} \cdot \overleftrightarrow{\pi}_{\parallel} - \vec{R}] \right]. \end{aligned}$$

A more general gyrokinetic version is needed to include $k_{\perp} \rho_i \sim 1$ physics.

- Since by construction $\int d^3v (1, \vec{v}, v_r^2) F = 0$,

the kinetic distortion F does not produce any extraneous δn , $\delta \vec{V}$ or δp .

Hence it produces no $\delta \vec{J}$, which is consistent with extended MHD since

$\vec{J} = \vec{\nabla} \times \vec{B}$ is “owned” by the Maxwell and MHD equations.

- The \vec{q}_s , $\overleftrightarrow{\pi}_s$, \vec{R}_s , Q_s velocity-space closure and collisional moments

$$\vec{q} \equiv \int d^3v \vec{v} \left(\frac{mv^2}{2T} - \frac{5}{2} \right) \bar{F}, \quad \overleftrightarrow{\pi} \equiv \int d^3v m (\vec{v}\vec{v} - v^2 \mathbf{1}/3) \bar{F}, \quad \vec{R} \equiv \int d^3v m \vec{v} \mathcal{C}\{\bar{F}\}, \quad Q \equiv \int d^3v \frac{mv^2}{2} \mathcal{C}\{\bar{F}\}$$

will be consistent with extended MHD and transport equations.

²Z. Chang and J.D. Callen, “Unified fluid/kinetic description of plasma micronstabilities. Part I,” Phys. Fluids B 4, 1167 (1992).

³S. Bruner, E. Valeo, J.A. Krommes, “Collisional delta-f scheme with evolving background for transport time scale simuls,” Phys. Pl. 6, 4504 (1999).

⁴J.J. Ramos, Phys. Pl. 17, 082502 (2010); J.J. Ramos, Phys. Pl. 18, 102506 (2011); B.C. Lyons, S.C. Jardin, J.J. Ramos, Phys. Pl. 19, 082515 (2012).

There Can Be Inconsistency Issues With Gyrokinetics

- The gyrokinetic formalism, in its usual implementations, is inconsistent with extended MHD and transport equations because
 - it is based on the reduced MHD axisymmetric magnetic field representation which neglects compressional Alfvén wave constraints on V_{tor} and evolving \vec{B} ,
 - its “drive” terms are of the form $\vec{v}_d \cdot \vec{\nabla} f_{\text{Max}}$ which is not consistent with Chapman-Enskog-type drives [$\sim \vec{v} \cdot (nq\vec{E} - \vec{\nabla}p)$, $\vec{v} \cdot \vec{\nabla}T$, $(\vec{v}\vec{v} - (v^2/3)\mathbf{I}) : \mathbf{W}$];
 - thus, its “input” does not include many V_{tor} , 3-D $\delta\vec{B}$ effects and its “output fluxes” will be inconsistent with the tokamak plasma transport equations.
- Most gyrokinetic studies concentrate on microturbulence & fluxes caused by particle drifts across axisymmetric surfaces and they
 - employ a large $V_{\text{tor}} \simeq E_\rho/B_p$ approximation plus diamagnetic flow but neglect the ion poloidal flow in the radial force balance (V_{tor}) equation,
 - concentrate on high n modes and do not treat low n global modes well because such modes require $\vec{J} = \vec{\nabla} \times \vec{B} / \mu_0$ with $\partial\vec{B}/\partial t = \vec{\nabla} \times \vec{E}$ in which \vec{E} is from Ohm’s law plus there are dissipative layer effects at rational surfaces,
 - treat non-ambipolar transport fluxes inconsistently — magnetic flutter determines local E_ρ but many other non-ambipolar transport effects (e.g., \perp viscous, 3-D and NBI torques) are neglected in E_ρ determination.

Tokamak Plasma Transport Equations Include Many Effects

- Transport equations for n_e , $L_{\text{tor}} \equiv \rho_m \langle R^2 \rangle \Omega_{\text{tor}}$, p_s , ψ_p with sources:⁵

$$\text{density} \quad \frac{1}{V'} \frac{\partial}{\partial t} \Big|_{\psi_p} n_e V' + \dot{\rho}_{\psi_p} \frac{\partial n_e}{\partial \rho} + \frac{1}{V'} \frac{\partial}{\partial \rho} (V' \Gamma) = \langle \bar{S}_n \rangle,$$

$$\text{toroidal momentum} \quad \frac{1}{V'} \frac{\partial}{\partial t} \Big|_{\psi_p} L_{\text{tor}} V' + \dot{\rho}_{\psi_p} \frac{\partial L_{\text{tor}}}{\partial \rho} + \frac{1}{V'} \frac{\partial}{\partial \rho} (V' \bar{\Pi}_{\rho\zeta}) = \langle \vec{e}_\zeta \cdot (\overline{\vec{J} \times \vec{B}} - \vec{\nabla} \cdot \overleftrightarrow{\bar{\Pi}} + \bar{S}_{\vec{p}}) \rangle,$$

$$\text{energy} \quad \frac{3}{2} p_s \frac{\partial}{\partial t} \Big|_{\psi_p} \ln p_s V'^{5/3} + \frac{3}{2} \dot{\rho}_{\psi_p} \frac{\partial p_s}{\partial \rho} + \frac{1}{V'} \frac{\partial}{\partial \rho} (V' \Upsilon_s) + \langle \vec{\nabla} \cdot \vec{q}_{*s}^{\text{pc}} \rangle = \bar{Q}_{\text{net } s},$$

$$\text{poloidal flux} \quad \frac{\partial \psi_p}{\partial t} \Big|_{\psi_t} = D_\eta \Delta^+ \psi_p - S_{\psi_p}, \quad D_\eta \equiv \frac{\eta_{\parallel}^{\text{nc}}}{\mu_0}, \quad S_{\psi_p} = \frac{\partial \Psi_p}{\partial t} + \frac{\eta_{\parallel}^{\text{nc}}}{I \langle R^{-2} \rangle} \langle \vec{B}_0 \cdot \vec{J}_{\text{drives}} \rangle.$$

- There are many classes of effects in plasma transport equations:
 - transients in the poloidal flux ψ_p via $\partial/\partial t|_{\psi_t}$ and advection of ψ_p surfaces relative to the toroidal-flux-based radial coordinate ρ via $\dot{\rho}_{\psi_p} \equiv \dot{\psi}_p/\psi'_p$,
 - transport fluxes of ambipolar density Γ , total momentum $\bar{\Pi}_{\rho\zeta}$ and heat Υ_s , \vec{q}_{*s}^{pc} “radially” across ψ_p poloidal flux surfaces that have many contributions induced by collisions, microturbulence and 3-D effects on each species s ,
 - toroidal ($\vec{e}_\zeta \equiv R^2 \vec{\nabla} \zeta$) plasma torques caused by $\overline{\vec{J} \times \vec{B}}$ and viscous stresses $\vec{\nabla} \cdot \overleftrightarrow{\bar{\Pi}}$,
 - ambipolar density $\langle \bar{S}_n \rangle$, toroidal momentum $\langle \vec{e}_\zeta \cdot \bar{S}_{\vec{p}} \rangle$ and energy $\langle \bar{S}_\mathcal{E} \rangle$ sources that contribute to the net energy heating rate $\bar{Q}_{\text{net } s}$, which includes Joule and external heating sources (NBI, ICRH, ECH etc.) plus radiation losses.

⁵J.D. Callen, C.C. Hegna, and A.J. Cole, “Transport equations in tokamak plasmas,” Phys. Plasmas **17**, 056113 (2010).

Plasma Toroidal Rotation Equation Includes 3-D Effects

- The magnetic field will be represented in ψ_p, θ, ζ coordinates by

$$\vec{B} = \underbrace{\vec{B}_0(\psi_p, \theta)}_{\text{2D, axisymm.}} + \sum_{n,m \neq 0} \underbrace{\delta \vec{B}_n(\psi_p, m) \cos(m\theta - n\zeta - \varphi_{m,n})}_{\text{low } m, n \text{ resonant, non-resonant}} + \underbrace{\delta \vec{B}_N(\psi_p, \theta) \cos(N\zeta)}_{\text{medium } n, \text{ ripple}} + \underbrace{\dots}_{\mu\text{turb.}}$$

- On μs time scale compressional Alfvén waves enforce radial force balance:

$$\Omega_{\text{tor}} \equiv \vec{V}_i \cdot \vec{\nabla} \zeta = - \left(\frac{\partial \Phi_0}{\partial \psi_p} + \frac{1}{n_i q_i} \frac{\partial p_i}{\partial \psi_p} \right) + q \vec{V}_i \cdot \vec{\nabla} \theta \quad \Longrightarrow \quad V_{\text{tor}} \simeq \frac{E_\rho}{B_p} - \frac{1}{n_i q_i B_p} \frac{dp_i}{d\rho} + \frac{B_t}{B_p} V_{\text{pol}i}$$

- On the ms time scale poloidal flow is damped to $V_{\text{pol}i}^{\text{nc}} \simeq (c_p/q_i)(dT_i/d\psi_p)$.
- Toroidal plasma torques cause radial particle fluxes: $\vec{e}_\zeta \cdot \vec{F}_{\text{orce}} = -q_s \vec{\Gamma}_s \cdot \vec{\nabla} \psi_p$.
- Setting the total radial plasma current induced by sum of the non-ambipolar particle fluxes to zero yields transport equation^{5,6} for plasma toroidal angular momentum density $L_{\text{tor}} \equiv \sum_{\text{ions}} m_i n_i \langle R^2 \vec{V}_i \cdot \vec{\nabla} \zeta \rangle$, $\Omega_{\text{tor}}(\rho, t) \equiv L_{\text{tor}}/m_i n_i \langle R^2 \rangle$:

$$\underbrace{\frac{\partial L_{\text{tor}}}{\partial t}}_{\text{inertia}} \simeq - \underbrace{\langle \vec{e}_\zeta \cdot \vec{\nabla} \cdot \vec{\pi}_{i\parallel}^{\leftrightarrow 3D} \rangle}_{\text{NTV from } \delta B} + \underbrace{\langle \vec{e}_\zeta \cdot \delta \vec{J} \times \delta \vec{B} \rangle}_{\text{resonant FEs}} - \underbrace{\langle \vec{e}_\zeta \cdot \vec{\nabla} \cdot \vec{\pi}_{i\perp}^{\leftrightarrow} \rangle}_{\text{cl, neo, paleo}} - \underbrace{\frac{1}{V'} \frac{\partial}{\partial \rho} (V' \Pi_{i\rho\zeta})}_{\text{Reynolds stress}^6} + \underbrace{\langle \vec{e}_\zeta \cdot \sum_s \vec{S}_{ps} \rangle}_{\text{mom. sources}}$$

- Radial electric field for net ambipolar transport is determined by Ω_{tor} (from L_{tor}):

$$E_\rho \equiv -|\vec{\nabla} \rho| \partial \Phi_0 / \partial \rho \simeq |\vec{\nabla} \rho| [\Omega_{\text{tor}} \psi_p' + (1/n_{i0} q_i) dp_i/d\rho - (c_p/q_i) dT_i/d\rho], \quad \psi_p' \simeq R B_p$$

⁶J.D. Callen, A.J. Cole, C.C. Hegna, "Toroidal flow and radial paricle flux in tokamak plasmas," Phys. Pl. **16**, 082504 (2009); Err. **20**, 069901 (2013) .

3-D Field Effects In Toroidal Momentum Equation

- General transport equation for toroidal angular momentum density $L_{\text{tor}} \equiv \rho_m \langle R^2 \rangle \Omega_{\text{tor}}$ is⁶

$$\text{toroidal momentum} \quad \frac{1}{V'} \frac{\partial}{\partial t} \Big|_{\psi_p} L_{\text{tor}} V' + \dot{\rho}_{\psi_p} \frac{\partial L_{\text{tor}}}{\partial \rho} + \frac{1}{V'} \frac{\partial}{\partial \rho} (V' \overline{\Pi}_{\rho\zeta}) = \langle \vec{e}_\zeta \cdot \left(\overline{\vec{J} \times \vec{B}} - \vec{\nabla} \cdot \overleftrightarrow{\Pi} + \overline{\vec{S}_p} \right) \rangle,$$

- Small 3-D field ($|\delta \vec{B}|/B_0 \sim \varrho_*$) effects come about in many ways:⁷
 - externally applied resonant $m/n \simeq q$ and non-resonant fields cause field error (FE, $\langle \vec{e}_\zeta \cdot \overline{\vec{J} \times \vec{B}} \rangle$) and neo. toroidal viscous (NTV, $\langle \vec{e}_\zeta \cdot \vec{\nabla} \cdot \overleftrightarrow{\Pi}_{\parallel} \rangle$) damping of Ω_{tor} ,
 - toroidal magnetic field “ripple” $\hat{b}_0 \cdot \delta \vec{B}_N$ caused by the finite number of coils that produce the toroidal magnetic field, which damps Ω_{tor} via NTV,
 - externally applied edge resonant magnetic perturbations (RMPs) used to modify the pressure profile there and stabilize edge MHD instabilities, and
 - spontaneous magnetic perturbations in the plasma which are caused by extended MHD macroscopic plasma instabilities that are controlled, e.g., neoclassical tearing modes (NTMs) or resistive wall modes (RWMs).

⁷J.D. Callen, topical review on “Effects of 3D magnetic perturbations on toroidal plasmas,” Nucl. Fusion **51**, 094026 (2013).

Differences From Present Approaches Are Suggested

- The fluid moment equations should be accepted as the basis for self-consistent and comprehensive modeling of tokamak plasmas.
- Extended MHD should be used to ensure macroscopic stability, and provide ρ_* order flows in flow-shifted Maxwellian equilibrium plus evolving \vec{B} field for Chapman-Enskog-based kinetic analysis.
- Gyrokinetic community should accept that its role is not to “do everything” but use flows and \vec{B} from extended MHD as “input” and produce closure moments for transport equations as “output.”
- The tokamak plasma transport equations for electron density n_e , plasma toroidal rotation frequency $\Omega_{\text{tor}}(E_\rho)$, species pressures p_s and poloidal flux ψ_p should all be solved for simultaneously, i.e., $\Omega_{\text{tor}}(E_\rho)$ should not just be obtained from experimental data.
- This modular approach needs to be iterated in order to obtain self-consistency between the MHD, kinetic and transport models.

SUMMARY: Need Fluid-Based Modular Approach

- Plasma kinetics is fundamental, but for time scales longer than species collision times (i.e., $t > 1/\nu_s$) 3-D (\vec{x}) fluid equations (extended MHD and transport) are feasible, appropriate, useful and needed. They are exact to extent that the relevant Chapman-Enskog kinetic equation (CEKE) can be solved for the kinetic distortion F_s which yields needed closure and collisional moments.
- Integrated simulations of tokamak plasmas need modular approach:
 - use small gyroradius expansion to order tokamak time scales, especially in the radial, parallel, and toroidal components of the plasma force balances,
 - use extended MHD to check macrostability, and solve for order ρ_* flows plus evolving \vec{B} field with reconnection regions and plasma responses to 3-D fields,
 - solve relevant CEKE in flowing equilibrium and evolving \vec{B} field to obtain collision-, fluctuation- and 3-D- induced closures and transport fluxes,
 - solve tokamak plasma transport equations for n_e , $\Omega_{\text{tor}}(E_\rho)$, p_e , p_i , ψ_p ,
 - and iterate MHD, CEKE, closures and transport steps for self-consistency.